



Astrophysics with EMRIS, IMRIS, and XMRIS

Formation, evolution, and the information encoded in the gravitational wave signatures

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Apr. 3, 2025

Ten Years to LISA @Jet Propulsion Laboratory

Inspiraling systems: EMRIs, IMRIs, XMRIs

Extreme mass-ratio inspirals (EMRIs): $q = M/m \sim 10^5 - 10^7$

Super Massive Black Hole (SMBH) + Stellar mass black holes (BH) /
Neutron stars (NS)/ White dwarfs (WD)

e.g. Rubbo+2006; Amaro-Seoane+2007; Berry+2019

Intermediate mass-ratio inspirals (IMRIs): $q \sim 10^1 - 10^5$

Intermediate mass black hole (IMBH) + SMBH
or
IMBH + BH

e.g. Coleman+2002,Gair+2011; Konstantinidis+ 2013; Haster+2016; Amaro-Seoane+2018

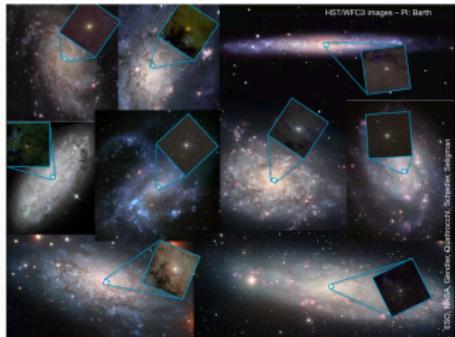
Extremely large mass-ratio inspirals (XMRIs): $q \sim 10^8$

SMBH + Brown dwarf (BD)

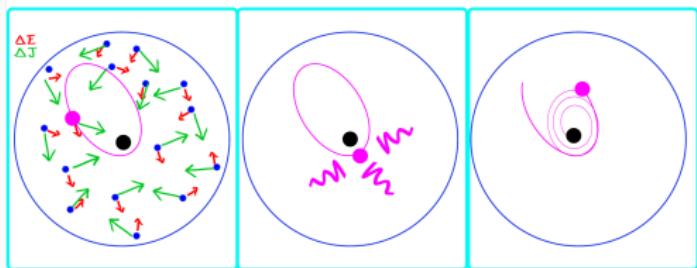
Amaro-Seoane,2019

Formation environment and mechanisms

Where are inspirals formed?



Main formation process: 2-body relaxation



The orbit reaches a critical semimajor axis a_{crit} such that

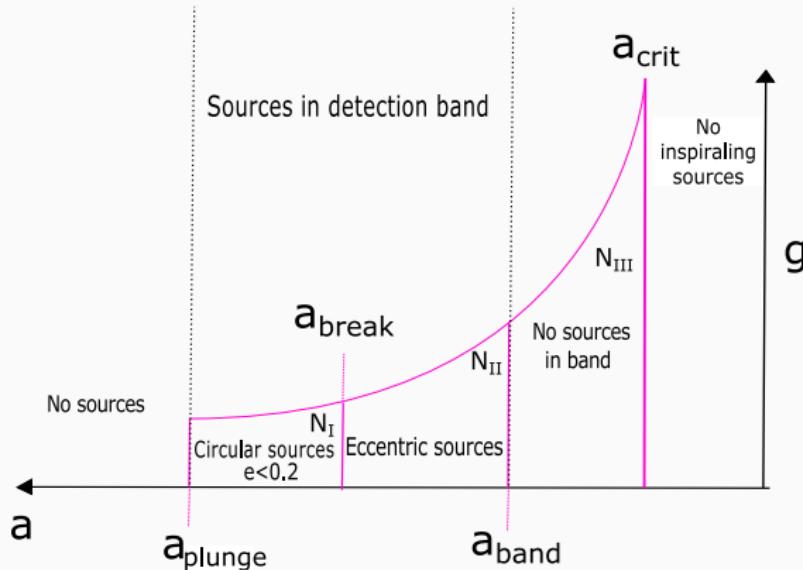
$$T_{\text{GW}} \lesssim T_{\text{rlx}}(a_0) \times (1 - e_0^2). \quad (1)$$

Hopman & Alexander 2005, Amaro-Seoane+2007

Galactic Nuclei and
Globular Clusters where
density can exceed
 $10^6 M_\odot pc^{-3}$

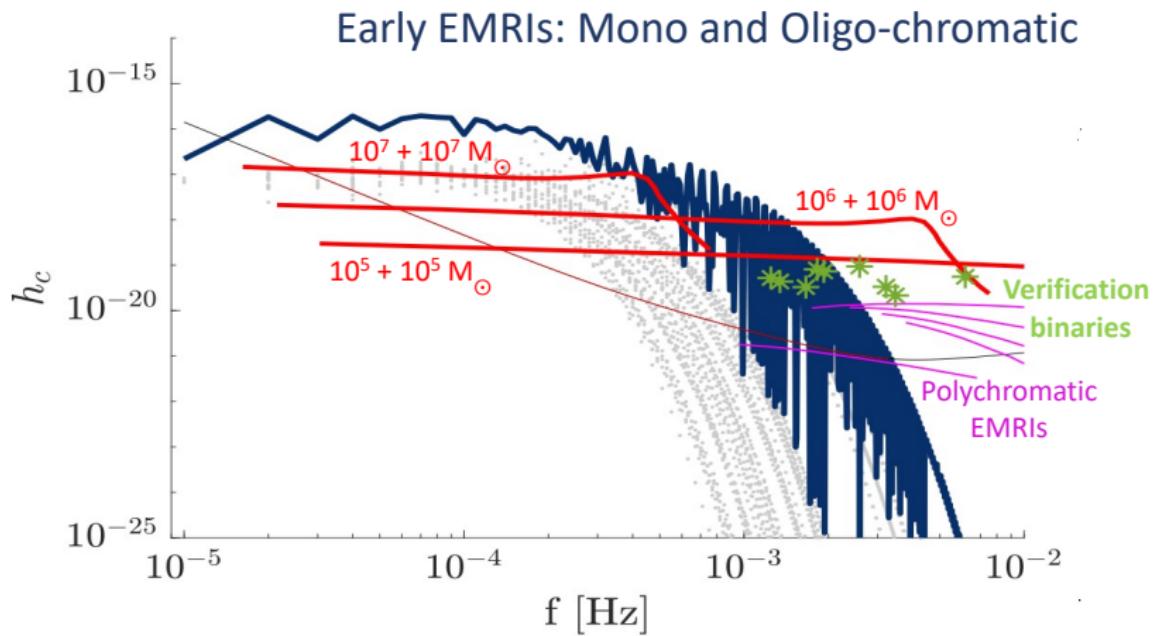
Other channels: Few body interactions

Number of detectable sources



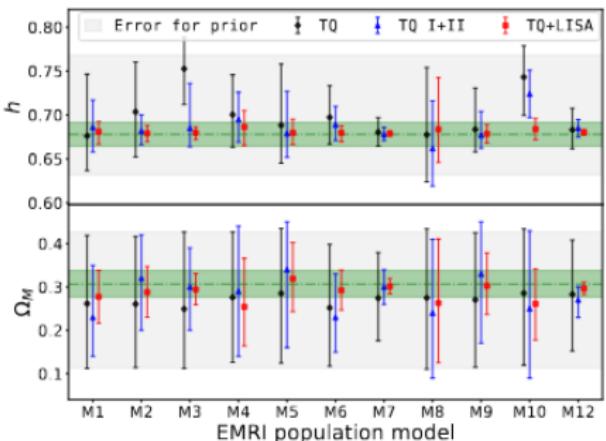
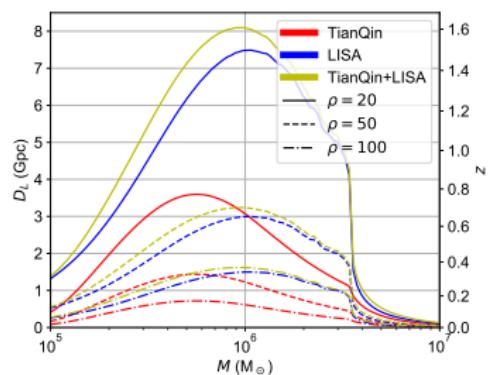
Amaro-Seoane+2019

Detectable sources: EMRIs



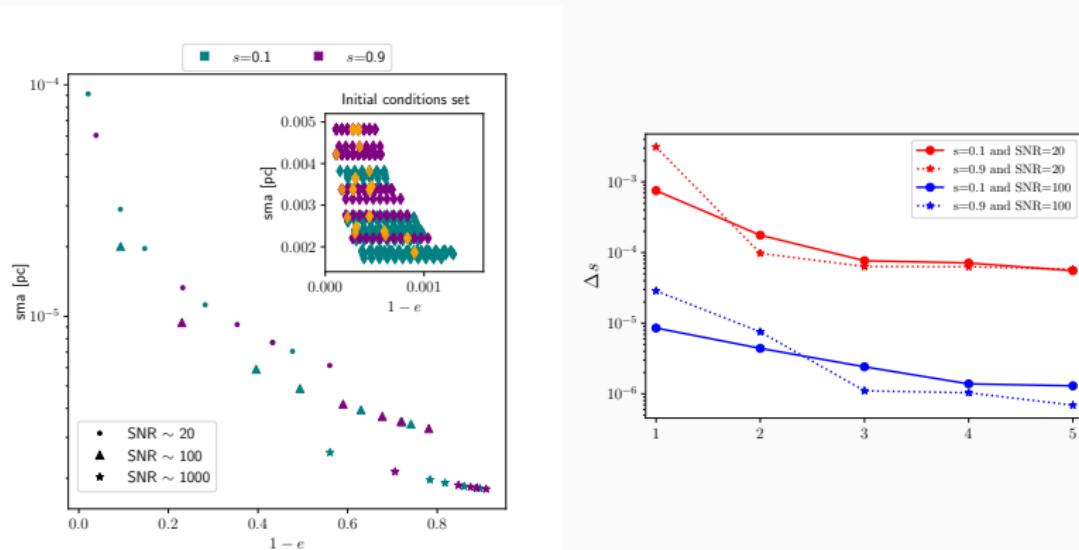
The Information Encoded in the Gravitational Wave Signatures

Cosmology with EMRIs



- EMRIs detectable to cosmological redshifts of $\lesssim 1.6$
- use as dark standard sirens to constrain Hubble constant $\delta h \sim 1\%$ and matter density $\delta \Omega_M \sim 10\%$

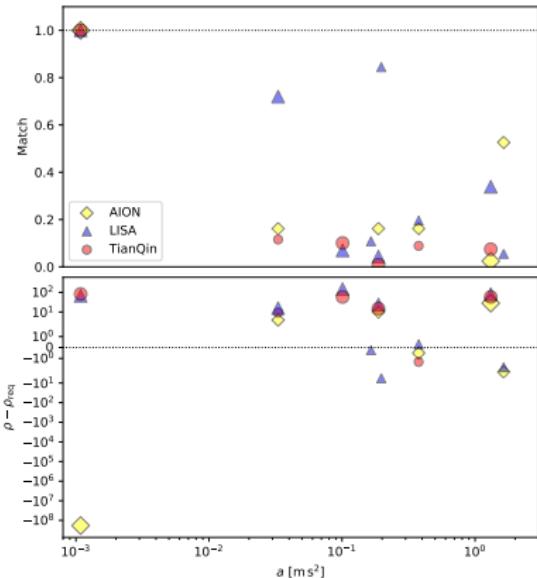
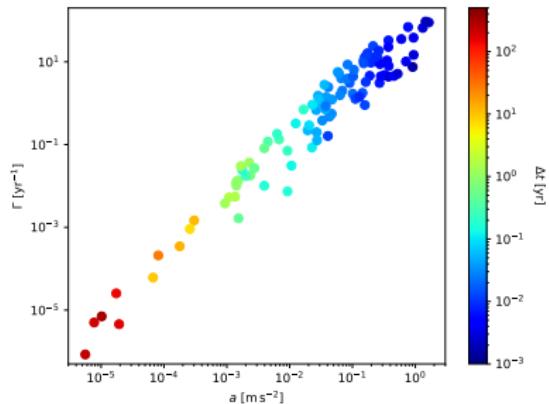
Measuring SgrA* with XMRIs



- multiple detectable XMRIs around SgrA*
- measure mass and spin of SgrA*: $\Delta M \sim 10^{-2} M_{\odot}$ and $\Delta s \sim 10^{-5}$

VVA+2023, VVA+2024

Studying Globular Clusters with IMRIs



- IMRIs undergo frequent weak encounters
- acceleration induces phase shift leading to significant mismatch

ATO+2025